Student Engineering Projects 2018



• Home

You are currently viewing an ING Confluence page.



Click on the ING icon to go to the ING homepage.

Astronomy



Links

- Applying for Time
- Archive
- Computing
- Conferences and Meetings
- o Detectors
- Exposure Time Calculator
- Fault Reports
- Feedback form
- o Filters
- o Instruments
- Library
- Object visibility
- Observing logs
- Override programmes
- Planning observations
- Publications
- Schedules
- Seminar programme
- Service programme
- Site quality
- Studentship programme
- Time allocations
- Weather and seeing

Telescopes



- William Herschel Telescope
- Isaac Newton Telescope
- Projects

These are ING's major projects:



Development of the WEAVE multi-object spectrograph. Replacement of the low-level Telescope Control System for the WHT. HARPS at the INT.

Other Links



Other Links

- Contacts
- Job opportunities
- Latest news

Public Information

- Ask an Astronomer
- Credits
- Images
- WHT Information
- INT Information
- News
- Photos and Videos
- Scientific Highlights
- News
- Virtual Tour
- Visiting the ING
- About the ING



Links

- Address and contacts
- Job opportunities
- Latest news
- Overview
- Staff
- Strategy

[INGPUBLIC]

This is an ING mailing list aimed at disseminating important research, development or outreach news, and information on public events organised by the Isaac Newton Group of Telescopes (ING), Roque de los Muchachos Observatory, La Palma. Messages sent to [INGPUBLIC] are also posted on the: ING home page ING RSS feed ING Twitter

Every posted item is moved to the previous news and events archive after some time. To subscribe or manage your subscription, visit the [IN GPUBLIC] home page.

[INGNEWS]

This is an important source of breaking news concerning the Isaac Newton Group of Telescopes (ING) especially with regard to telescope time. To subscribe or manage your subscription, visit: [INGNEWS] home page.

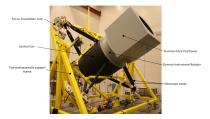
Student Engineering Projects for 2018 WEAVE

WEAVE is the next observing facility being built for the 4.2-m William Hershel Telescope. The facility comprises a new 2-degree field-of-view prime focus corrector with a pick-and-place fibre positioner, a small number of individually deployable integral field units (IFUs), and a large integral field unit. The IFUs and the MOS fibres can be used to feed a dual-beam spectrograph and the facility is expected to be on-sky by 2019.

Project 1: Focus Translation System PLC Modifications

Background

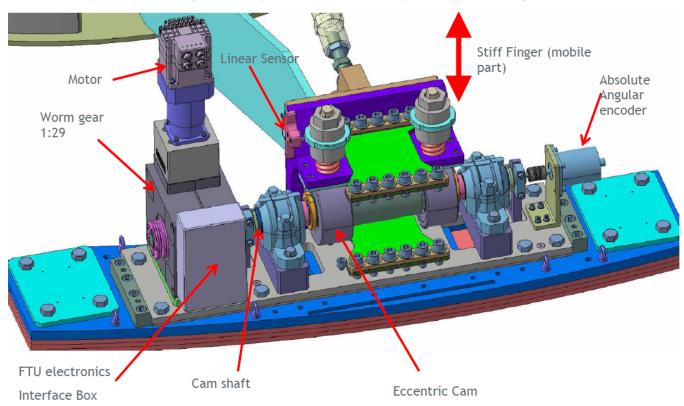
The new 2-degree field-of-view prime focus corrector, with a Fibre Positioner, will be placed within the top-end assembly of the telescope as shown in the following figure:



The WEAVE top-end assembly (WTA) consists of a mechanical representation of the Fibre Positioner and the Instrument Rotator which is connected to the actual Central Can that will house the Prime Focus Corrector. The WTA will be connected to the telescope top-end ring via eight telescope vanes connected to four Focus Translation Units (FTU); collectively known as the Focus Translation System (FTS). The whole of the top end is designed to move perpendicularly to the plane defined by the four FTUs. The permitted range of movement is small but allows the top end to be refocussed to deal with changes in the position of the corrector optics due to temperature changes and mechanical flexures induced in the structure as the telescope moves in elevation.

Situation

The FTU Control System, which was designed and built by SENER, uses an Allen Bradley PLC to send command signals to the FTU motors:



The individual demand values are provided by the Telescope Control System (TCS); one value for each motor. This means that for each movement of the FTS, four demand values are required to be sent from the TCS to the FTS Control System (FCS). This system works well but places an additional overhead on the TCS, the overhead is primarily concerned with calculating the individual demand values for the motors.

Project

Given that the available processing power of the PLC is greatly underused, it makes sense to move some of the logic, required to calculate the individual values, from the TCS to the FCS. This project will involve the design and delivery of the PLC code required to accept a single focus demand from the TCS, calculate the demand values for the individual motors and apply the required drive signals to the FTS drive system. In addition to this, a similar process will need to be implemented for the application of the FTS tilt as function telescope elevation.

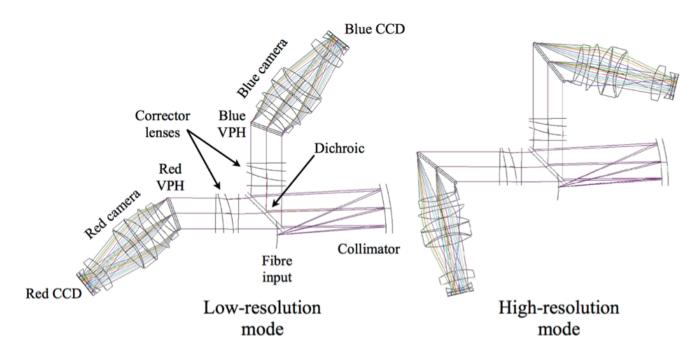
The application is designed using a client-server architecture. The server side of the application consists of a servlet running under the Apache Tomcat web server and will be developed in Java using the Google Web Toolkit framework (GWT) which is already used with other systems at the WHT. The client side of the application consists of a set of forms running on a web browser and allows the user to interact with the application. It will

be developed using Java and later converted into JavaScript by the GWT compiler so that it can run from any web browser. The purpose of this project is to write the code required to convert the design into reality.

Project 2: Planning integration of the spectrograph into GHRIL

Background

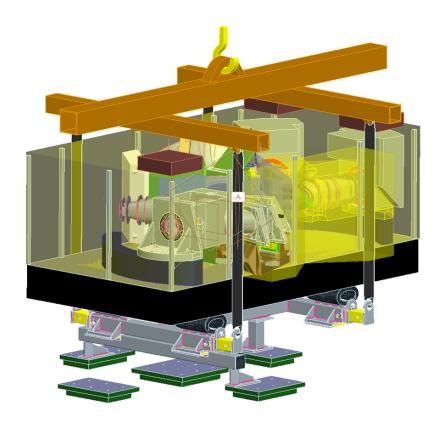
The WEAVE spectrograph is a dual-beam optical system which incorporates a dichroic at ~600nm and has a grating exchange mechanism that provides two resolution modes in each arm; a low-resolution mode at 5000 and a high-resolution mode at 20 000.



The WEAVE spectrograph will be permanently housed in the GHRIL enclosure (one of the two Nasmyth platforms of the WHT telescope), providing a stable and clean environment for the instrument. To better appreciate the size of the spectrograph, below is a picture of the collimator being prepared for polishing:



The spectrograph will be assembled on its optic table, enclosed in a light-tight assembly and the whole structure lifted into GHRIL:



Situation

The WEAVE spectrograph is composed of several pre-integrated sub-systems that will be assembled together at the WHT. The institution responsible for the development of the instrument will provide general guidelines and recommendations that will need to be further developed and detailed for the WHT and the working environment.

The current GHRIL enclosure will be refurbished to be turned into a clean room where dust, temperature and humidity will be monitored and controlled.

Project

ING will manage the integration process from the delivery of the crates to the WHT and the installation of the spectrograph into the GHRIL enclosure. The ING will also provide support for the assembly of the instrument in GHRIL and will ensure that the clean room environment is maintained.

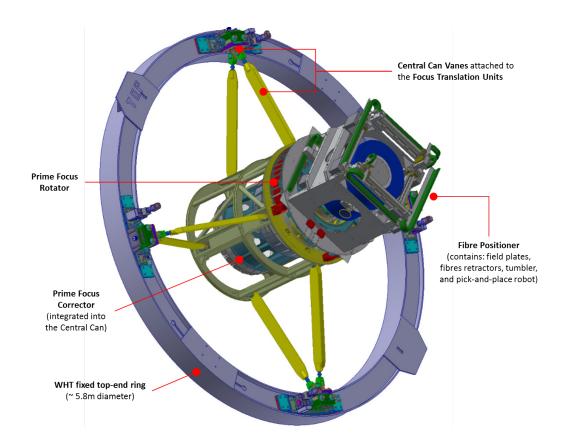
The complete process for integrating the spectrograph into GHRIL needs to be designed and detailed as far as possible, in order to:

- 1. Allow early procurement of any specific tools that may be needed
- 2. Anticipate any potential issues and reduce risks
- 3. Estimate the resources that will be required to deliver the working environment

Project 3: Planning the optical assembly and testing of the corrector

Background

The Prime Focus Corrector is a sub-system of the WEAVE Focal Plan Assembly (WFPA), interfacing with the Focus Translation System (as described in Project 1), the Positioner System and the optical fibres feeding the spectrograph. The corrector is made of a mechanical barrel that contains six corrector lenses assembled in their cell and aligned with respect to the main optical axis of the telescope.



The purpose of the WEAVE Prime Focus Corrector is to correct for the optical aberrations in the two-degree field-of-view. The corrector also includes an Atmospheric Dispersion Corrector (ADC) which consists of two pairs of counter-rotating, air-separated doublets, to compensate for atmospheric dispersion whilst the telescope is moving in elevation.

Situation

The corrector is currently being manufactured by SENER and will be delivered as a complete, fully-aligned unit. The Focus Translation System is built and includes mechanical dummies which represent the corrector and rotator. This will be used to simulate the volume and mass of the final assembly.

The philosophy for assembling the corrector has already been drafted and is split in two distinct phases:

- 1. The corrector will be integrated within the Central Can and aligned with respect to the rotator mechanical axis of rotation.
- 2. The whole WFPA (with the corrector) will be mounted onto the telescope top-end and will be aligned with respect to the WHT optical axis (defined by the primary mirror).

Phase 1 will be performed while the WFPA is on its trolley situated at the WHT ground floor. It will involve the opto-mechanical alignment as well as testing of the control software (to rotate the ADC mechanisms).

Phase 2 will require the telescope optical axis to be identified. The various targets that will be used will be previously positioned to ensure an absolute reference is set.

Fine tuning of the WFPA alignment will be carried out on sky by pointing at bright stars.

Project

The project will involve work to deliver Phases 1 and 2. The corrector will be optically tested to ensure the optical performance is met. For this, various tests will be designed and planned into the main assembly philosophy. The opto-mechanical tools and alignment jigs will be specified, procured, and tested.

Project 4: WebSocket-based communications framework for GUIs

The ING is investigating a possibility of moving most of the user interfaces to the browser. This has many benefits compared to classical desktop applications and should make deployment of applications much easier. In the past, web-based GUIs with a live content use the so-called AJAX communication framework. A disadvantage of this is that it has a significant overhead when transferring data and can thus cause more load on the network infrastructure and can slow down the applications. A more efficient method, called the WebSocket protocol, is now supported by all major browsers. WebSockets are rather low-level, and can make programming tedious. A higher-level protocol which builds on top of WebSockets is called WAMP and promises to provide mechanisms such as publisher-subscriber to web-based applications. Although this protocol is very new and still under development, there are already implementations for many programming languages.

The ING has carried out some experiments using raw WebSockets, but has no experience with WAMP.

This project consists of creating demo versions of servers using languages used at the ING (Python, Java, C++) and the client (GUI) versions using preferably Dart or Javascript. The task is to provide patterns for the creation of servers and clients which would be easily reused by the ING engineers when creating new web-based applications.

Project 5: WEAVE sequencer

Before each observation or calibration with the WHT is carried out, the instruments need to be properly configured with a sequence of commands, by taking exposures. Due to the large number of mechanisms, light sources and other instrumentation, the configuration process may become rather complicated. The problem may be addressed by writing scripts, but this method lacks flexibility and does not allow easy monitoring of the status of the sequence and of the configured mechanisms. The ING has therefore developed a framework for developing sequences. The sequences consist of steps and steps consist of actions, which are designed to issue a single command to instrument controllers. After each action, a series of predetermined checks are automatically performed and the outcome of each check is shown in the GUI. If the action or the check fails, the sequence is stopped and the user can choose between different options to proceed such as aborting the sequence, skipping the action or continuing.

An important feature of the sequencer is the ability to build sequences without having the actual code. This allows users, who are not programmers, to build sequences without having to deal with programming details.

All sequences and the underlying code are stored in a database. As the foundation of the sequencer grows, there will be less code to write and even the non-programmers will be able to write fully functioning sequences.

The sequencer is an existing application, which has been developed at the ING. It will be used for the WEAVE project and also for observation and calibration sequences on other instruments.

Project

The current code base needs enhancing. Some parts of code may need to be updated to use the current Python capabilities. Certain functional upgrades are needed such as integrated version control. Also, an intensive stress and performance testing plan will be required.

The application is written in Python 3 and uses wxPython GUI library and the PostgreSQL database. A good understanding of concurrent processing techniques is required for this project.